# Recollimation Shock, Transverse Waves and the Whip in BL Lacertae

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The MOJAVE Program is supported NASA Fermi Grant NNX08AV67G

# Monitoring Of Jets in Active Galaxies with xperiments

## Very Long Baseline Array



## **BL Lac Topics**

- Components, Ridge Line
- Recollimation Shock
- PA Variations: Wobble
- Transverse Waves
- Wave Speeds
- Relaxation in 2010: Wiggle
- Conclusions

## Components

- Components are a small number of elliptical Gaussians that sum to the image. They usually are circular.
- A component is often a bright spot in the image.
- Components are tracked in time, if the cadence of observations is fast enough.

## Ridge Line

- BL Lac is elongated and has a ridge.
- Most components move downstream
- along the ridge (±0.1 mas)

## **Ridge Line and Components**



## Several New Components per Year Max Speed 10.6c





## **Recollimation Shock**

- Component 7 = recollimation shock
- Analogy to M87, 3C120
- Simulation (Lind et al 1999, ...) shows a 'magnetic chamber' and fast cpts ejected into a 'nose cone' (not full 3D RMHD) (Meier p715)
- Need strong toroidal component in the magnetic field

## Recollimation Shock II

		Z	vmax	dist from core
•	3C 120	0.033	5.3c 3.1 *	~ 5 mas 80 mas ~ 3x10 <sup>7</sup> r <sub>a</sub>
•	M 87 BL Lac	.00436 .0686	4.3 ** 10.6 #	>= $10^{6} r_{g}$ ~ $10^{6} r_{g}$
<ul> <li>* downstream from C1</li> <li>** downstream from HST1</li> <li># downstream from cpt 7</li> <li>MOJAVE</li> </ul>				

BL Lac Position Angle vs Epoch



#### Transverse Wave 2004 - 2007



## Shifted Ridge Lines



#### Transverse Waves 1999-2000



### Transverse Wave 2000-2001



## Component 16 Advected with the Transverse Wave



## Several New Components per Year Max Speed 10.6c





# BL Lac 2010.8-2012.8 wiggle





## BL Lac Conclusions I

- Strong quasi-stationary component is identified as a recollimation shock.
- Distance from core ~  $10^6 r_g$
- Superluminal components appear to come from or through the recollimation shock.

## BL Lac Conclusions II

- Jet supports transverse waves.
- Waves have superluminal speed.
- Waves are correlated with swings in the inner PA.
- Components are advected with the transverse motion.
- Jet acts more like a rope than a water hose.
- When wave activity dies down, a stable wiggle appears.

## Conclusions III

These observations support a model in which the jet contains a strong toroidal magnetic field. Transverse waves (Alfven?) are excited by PA swings of the nozzle, and propagate superluminally downstream, as largescale wiggles on the ridge line. The superluminal components (fast MHD waves?) stay on the ridge, and can be advected transversely. In 2009 the waves died down, and a small-scale stationary wiggle appeared on the jet.